





## Numerical Experiments with Shock Turbulence Interaction

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## Motivation (1)



Shock-turbulence interaction is central to high-speed propulsion systems yet its current models are quite inaccurate

Examples: scramjet engine; supersonic inlets

- Control of shockwave-turbulent boundary interactions in supersonic aircraft inlets (bleed system) causes a 15-20% performance penalty;
- Stability of the shock-train system limits the operability of a scramjet engine



Shock-Boundary layer interaction

#### Scramjet engine





- Interactions between strong shocks and interfaces lead to flow instabilities, which drive turbulent mixing
- Accurate prediction of shock-induced turbulent mixing remains an open problem
  - Examples: Rayleigh-Taylor mixing, Richtmyer-Meshkov mixing
  - Applications: inertial confinement fusion, supernova explosions



Source: lasers.llnl.gov/programs/nic/icf/

Inertial confinement fusion by shock-induced implosion of a deuterium/tritium capsule



Source: flash.uchicago.edu

Supernova explosion

## The key challenge:



- Simulation of broadband turbulence requires numerical methods with minimum numerical dissipation
  - Capture the widest range of resolvable scales on a given grid
  - Spectral methods; high-order methods
- Simulation of flows with shocks requires numerical dissipation for shock capturing
- The key challenge:

How to design a numerical method that handles these conflicting requirements?

# Assessment of Numerical Algorithms: (for shock-turbulence interaction simulations)



- Broad range of high-resolution algorithms
  - High-order WENO (7th order)
  - High-order compact differences with selective artificial viscosity (STAN; STAN-I) Cook & Cabot, 2007, and modified Kawai & Lele, 2008; Mani, Larsson & Moin, 2009, Bhagatwala & Lele, 2009
  - Hybrid (8th order central differening + WENO) Larsson & Gustavsson, 2008
  - High order central difference + Wavelet sensor based artificial dissipation ADPDIS3D (Yee & Sjogreen, 2007)
- Broad range of benchmark test cases
  - Noh problem (strong implosion)
  - Shu-Osher problem
  - Interaction of vorticity/Entropy wave with normal shock
  - Taylor-Green problem
  - Compressible isotropic turbulence

Details in paper by Johnsen et al. to appear in J. Comput. Phys, 2009

## Numerical Algorithm Assessment: Compressible Isotropic Turbulence (with eddy shocklets)





Code	Color	Line style
Reference	black	varying
Stan	red	dashed
Stan-I	magenta	dashed (thin)
Hybrid	blue	solid
WENO	$_{ m cyan}$	solid (thin)
ADPDIS3D	green	dashed-dotted
Shock Fit	black	dotted

Table 2: Color and line legend for the plots.

Figure 11: Temporal evolution of the variance of different quantities for the isotropic turbulence problem on  $64^3$  grid. The reference is the solution on a 256<sup>3</sup> grid spectrally filtered to a  $64^3$  grid (circles).

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Figure 13: Spectra at  $t/\tau = 4$  for the isotropic turbulence problem on a  $64^3$  grid. The reference is the solution on a  $256^3$  grid.

### Numerical Algorithm Assessment: Effective Bandwidth

 Use compensated spectra E(k)/E<sub>exact</sub> (k) to define effective bandwidth





Figure 15: Compensated spectra on a  $64^3$  grid. As a reminder, *Stan* is dashed red, *Stan-I* is dashed magenta (thin), *Hybrid* is solid blue, *WENO* is solid cyan (thin), and *ADPDIS3D* is dash-dotted green.

## Canonical shock/turbulence interaction



- Isotropic turbulence passing through a normal shock in a perfect gas
- Isolates the core interaction between turbulence and shock



Eddies visualized by Qcriterion, colored by vorticity. Shock visualized by dilatation contour. Some past work on canonical shock/turbulence interaction



- Theoretical
  - Ribner (NACA 1953, NACA 1954, AIAA J 1987), Moore (NACA 1954): linear interaction analysis including Rankine-Hugoniot relations
  - Lele (PhysFI 1992): turbulent shock relations using RDT, predicted modified mean shock jumps
- Experimental: wind tunnels
  - Barre et al (AIAA J 1996): grid turbulence passing through a Mach 3 shock, hot-wire and LDV measurements
- Experimental: shock tubes
  - Hesselink & Sturtevant (JFM 1988): found "peaked" and "rounded" instantaneous pressure profiles in random medium
- Computational (DNS):
  - Lee et al (JFM 1993): resolved viscous shock structure, M = 1.2,  $\text{Re}_{\lambda}$  = 20, M<sub>t</sub> = 0.1, found modified instantaneous profiles of dilatation
  - Lee et al (JFM 1997): captured shock up to M = 3, found good agreement with Ribner's linear theory
  - Mahesh et al (JFM 1997): influence of entropy fluctuations
  - Ducros et al (2000): LES; Jamme et al (FTC 2002)

Post-shock Kolmogorov length scale: Scaling Estimate



• From the (incompressible) definition of the Kolmogorov length scale

$$\eta = \left(\frac{\nu^3}{\varepsilon/\rho}\right)^{1/4} \sim \left(\frac{\mu^2}{\rho^2 \omega^2}\right)^{1/4} \sim \frac{T^{3/8}}{\rho^{1/2} \omega^{1/2}}$$

where we used

$$arepsilon pprox \mu \omega^2 \;,\;\; 
u = \mu / 
ho \;,\;\; \mu \sim T^{3/4}$$

• Across the shock, assume that the vorticity changes as the density ratio; gives

$$\frac{\eta_{\rm d}}{\eta_{\rm u}} \approx \left(\frac{T_{\rm d}}{T_{\rm u}}\right)^{3/8} \left(\frac{\rho_{\rm d}}{\rho_{\rm u}}\right)^{-1}$$

in terms of Rankine-Hugoniot relations

Post-shock Kolmogorov Scale determines the resolution requirement for turbulence





- Past studies:  $\text{Re}_{\lambda}$  = 20 on 129\*64<sup>2</sup> grids (Lee et al, 1993, 1997)
- Present:  $Re_{\lambda} = 40$  on  $1040^*384^2$  grids;  $Re_{\lambda} = 60$  on  $1675^*512^2$  grids
- Return to local (small-scale) isotropy in present DNS, but not in past work
- Vorticity redistribution is a nonlinear process
  - Past work did not fully capture this due to lack of grid resolution
  - Has led to persistent misunderstanding of vorticity evolution at shock



#### Anisotropic post-shock turbulence





Eddies visualized by Q-criterion, colored by angle between vorticity vector and x-axis. Shock visualized by dilatation contour.

#### Reynolds stresses: amplification across the shock vs Mach number



- Amplification of kinetic energy agrees very well with Ribner's linear theory (and Lee97)
  - Kinetic energy must be governed largely by linear processes
- DNS (and experiment) yields much larger anisotropy
  - Nonlinear redistribution processes must be important







Complex shock structure in nonlinear regime





- Shock no longer a simple object
  - Qualitative agreement with Zank et al (PhysFI 2002)



- Define shock-dilatation as  $\theta_{
  m shock}(y,z) = \min_{x} \theta(x,y,z)$
- Study instantaneous profiles through the shock at extreme values of shockdilatation:
  - "Strong" interactions: most compressive events
  - "Weak" interactions: least compressive events



PDF of shock dilatation

#### Unsteady shock motion





Shock visualized by dilatation contour, colored by streamwise momentum.

#### "Strong" events: over-compression and post-shock expansion



#### "Weak" events: smooth profiles in nonlinear (high M<sub>t</sub>) regime



Hesselink & Sturtevant (JFM 1988)



• "Peaked" and "rounded" pressure traces for shock propagating through a random mixture of He and R12 at M = 1.007 - 1.1







- Direct numerical simulations of canonical shock/turbulence interaction that truly resolves the post-shock turbulence
- Disagreements with linear theory and past computations:
  - Vorticity components return to isotropy
  - Qualitatively different, and much larger, stress anisotropy
- Agreement with linear theory and past computations:
  - Kinetic energy amplification -- governed by linear processes
- Instantaneous shock/turbulence interaction dynamics:
  - Shock weakens (to the point of disappearing!) locally -- smooth profiles through shock in nonlinear regime
  - "Strong" shock events with up to twice the average compression



A more complex flow with shock-turbulence interaction



Expt. By Santiago & Dutton, 1997, JPP; Ben Yakar et al. 2006, PoF LES by Kawai & Lele, AIAA 2009-3995

#### JISC with Turbulent Crossflow





#### Time-Averaged Flowfields







Turbulent kinetic energy



Jet fluid (passive scalar)





#### **3D Flow Structures**











- Revisit to canonical shock-turbulence interaction (DNS)
- Other studies (DNS/LES) of shock-turbulent BL interaction, compression ramp..
- More complex engineering flows involving injection, mixing and combustion in high-speed flows

These are good laboratories for improving our understanding of turbulence behavior, its engineering modeling (RANS & LES), and for studies of aeroacoustics and flow control.

What will the next 50 years bring ?

Exciting science and collaborative research for sure. Happy 50 ! And Thanks.



## Additional slides



- Effect of turbulent Mach number on amplification of turbulence -- viscous effects make conclusions difficult
  - New  $\text{Re}_{\lambda}$  = 60 runs should clarify this
- Why is peak shock-motion at larger length scales than peak turbulence energy?
  - Theory suggests shock-motion "enslaved" to incoming turbulence
  - Need runs with higher k<sub>0</sub> to allow for deeper analysis
- What is the relation between shock-motion and Reynolds number? Or is the shock-motion dependent on some unknown parameter that was not matched?
- What is the turbulence structure behind the shock -- statistics show elongation in streamwise direction, but what does it look like?
  - More, and more intelligent, visualization needed



- Interesting finding: spectrum of shock motion seems to scale with ratio of turbulent pressure fluctuation to shock-induced pressure jump:  $M_t^2 / (M^2-1)$
- Puzzle #1: turbulence has peak energy at k=4, but shock motion has peak at k=1-2
- Puzzle #2: thickness of averaged shock increases with Re



#### Mean profiles





• Turbulence modifies the Rankine-Hugoniot relations (Lele, PhysFI 1992):

$$\overline{\rho}\widetilde{u} = \text{const}$$

$$\overline{\rho}\widetilde{u}^2 + \overline{p} + \overline{\rho}\widetilde{u''u''} = \text{const}$$

$$\overline{\rho}\widetilde{h_0}\widetilde{u} + \overline{\rho}\widetilde{h_0''u''} = \text{const}$$

## Numerical Algorithm Assessment: Taylor-Green Problem



 Inviscid problem with strong vortex-stretching Kinetic energy conservation



Figure 1: Mean quantities for the Taylor-Green vortex on a 64<sup>3</sup> grid. The zero subscript denotes the initial value.

## Numerical Algorithm Assessment: Taylor-Green Problem



Kinetic Energy Spectrum

(a) Convergence of the reference solution using the Hybrid code (b) Comparison between the different schemes. The reference in standard mode (solid) and with eight-order accurate dissi- solution is that obtained on the 256<sup>3</sup> grid using the Hybrid pation (dashed), on 256<sup>3</sup> (black), 128<sup>3</sup> (blue), and 64<sup>3</sup> (cyan). code.

Figure 2: Velocity spectra for the Taylor-Green vortex on  $64^3$  grid at t = 5.

Direct numerical simulation (DNS)



- Solve the Navier-Stokes equations without modeling
- Numerical challenges:
  - Shock-capturing requires numerical dissipation to smear shock
  - Numerical dissipation kills turbulence

